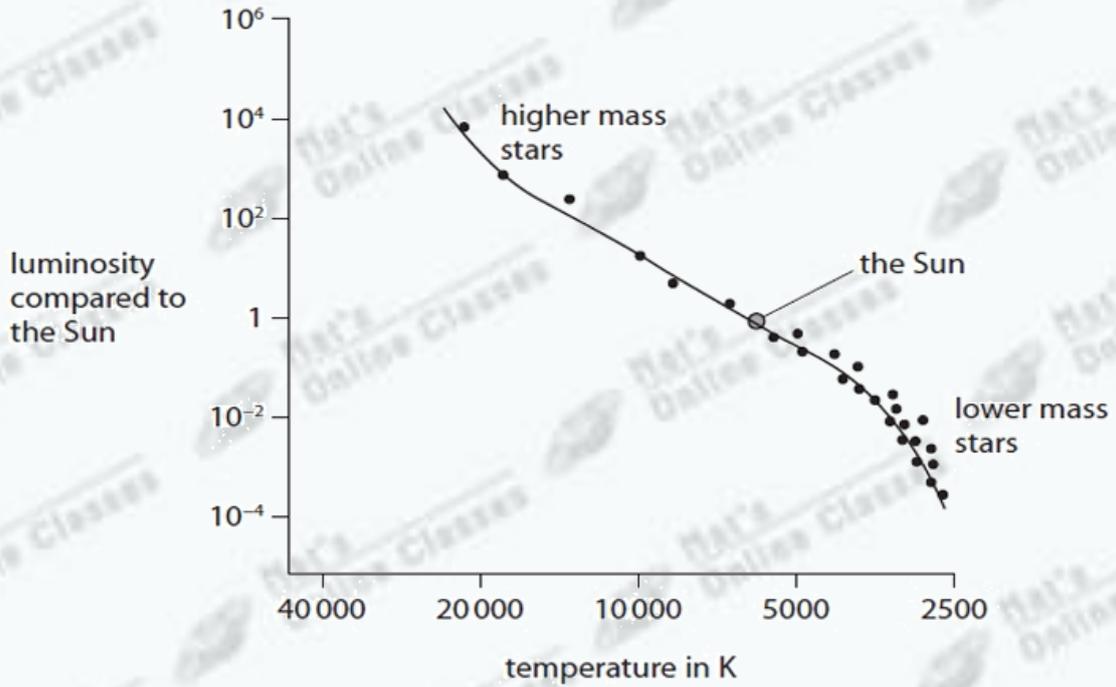


Stars

Total marks:13

Q1.

Figure 18 is a diagram giving some information about main sequence stars. Luminosity is a measure of how bright something is. An increase in luminosity means an increase in brightness.



© abyss.uoregon.edu

Figure 18

(i) Estimate the temperature of the Sun.

(1)

temperature of the Sun = K

(ii) State how the brightness of a main sequence star changes with its temperature.

(1)

.....
.....

(iii) State how the brightness of a main sequence star changes with its mass.

(1)

.....
.....

(Total for question = 3 marks)

Q2.

A nebula may evolve into a main sequence star, such as the Sun.

Explain how a nebula may evolve into a main sequence star.

(3)

.....
.....
.....
.....
.....

(Total for question = 3 marks)

Q3.

Stars may originate as a nebula.

(i) Describe the process that then occurs to produce the conditions necessary for nuclear fusion in a new star.

(3)

.....

.....

.....

.....

.....

.....

(ii) The energy, E , released in nuclear fusion is equivalent to loss in mass, m , according to the equation.

$$E = mc^2$$

where c is the velocity of light.

$$c = 3.00 \times 10^8 \text{ m/s}$$

In 1 second, the energy radiated by the Sun is $3.86 \times 10^{26} \text{ J}$.

Calculate the loss in mass of the Sun in 1 second.

(2)

loss in mass = kg

(Total for question = 5 marks)

Q4.

A star has evolved to become a neutron star.

The mass, M , of the neutron star, of radius R , is given by

$$M = \frac{4 \times \pi \times D \times R^3}{3}$$

where D is a constant

$$M = 4 \times 10^{30} \text{ kg}$$

$$D = 6 \times 10^{17} \text{ kg/m}^3$$

Use the equation to calculate the value for R .

(2)

$R = \dots\dots\dots \text{ m}$

(Total for question = 2 marks)